## Solution for Chapter 9

(compiled by Xinkai Wu)

1. Ex. 9.2 The holographically reconstructed wave [by Kip Thorne, Jaemo Park, and Alexei Dvoretskii/99]

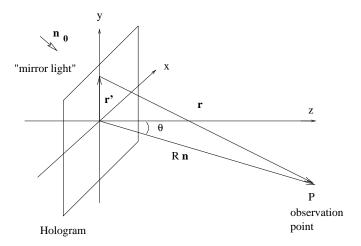


Figure 1: holographically reconstructed wave

(a) See Fig. 1, in which  $r=|\mathbf{r}|=|R\mathbf{n}-\mathbf{r}'|\approx R-\mathbf{n}\cdot\mathbf{r}'$ . Using the Helmholtz-Kirchhoff formula, we get

$$\psi_P = rac{ik}{2\pi} \int d{f S}' \left(rac{{f n}+{f n}_0}{2}
ight) rac{e^{i(kR-k{f r}'\cdot{f n})}}{r} \psi_0(x',y')$$

where

$$\psi_0(x',y') = T(x',y')e^{ik\mathbf{r}'\cdot\mathbf{n}_0}$$

$$\sim \left(|O|^2 + M^2 + MOe^{-ik\mathbf{r}'\cdot\mathbf{n}_0} + MO^*e^{ik\mathbf{r}'\cdot\mathbf{n}_0}\right)e^{ik\mathbf{r}'\cdot\mathbf{n}_0}$$

$$\sim (|O|^2 + M^2)e^{ik\mathbf{r}'\cdot\mathbf{n}_0} + MO + MO^*e^{2ik\mathbf{r}'\cdot\mathbf{n}_0}$$

and thus

$$\psi_{P} = \psi_{P}^{(1)} + \psi_{P}^{(2)} + \psi_{P}^{(3)}$$

$$\psi_{P}^{(1)} = \frac{ik}{2\pi} \int d\mathbf{S}' \left(\frac{\mathbf{n} + \mathbf{n}_{0}}{2}\right) (|O|^{2} + M^{2}) \frac{e^{ikR}}{r} e^{ik\mathbf{r}' \cdot (\mathbf{n}_{0} - \mathbf{n})}$$

$$\psi_{P}^{(2)} = \frac{ik}{2\pi} \int d\mathbf{S}' \left(\frac{\mathbf{n} + \mathbf{n}_{0}}{2}\right) MO \frac{e^{ikR}}{r} e^{-ik\mathbf{r}' \cdot \mathbf{n}}$$

$$\psi_{P}^{(3)} = \frac{ik}{2\pi} \int d\mathbf{S}' \left(\frac{\mathbf{n} + \mathbf{n}_{0}}{2}\right) MO^{*} \frac{e^{ikR}}{r} e^{ik\mathbf{r}' \cdot (2\mathbf{n}_{0} - \mathbf{n})}$$

If the mirror wave had been absent and the photographic plate replaced by a window, then by using Helmholtz-Kirchhoff with  $\tilde{\psi}_0 = O$  we would have gotten  $\psi_P = \frac{ik}{2\pi} \int d\mathbf{S}' \left(\frac{\mathbf{n}+\mathbf{n}_0}{2}\right) O^{\frac{e^{ikR-ik\mathbf{r}'\cdot\mathbf{n}}}{r}}$  which is the same as  $\psi_P^{(2)}$  (to within a multiplicative constant).

The direction of propagation for different terms can be obtained by finding the stationary phase points for rapidly oscillating terms (note the last factor in the expressions for  $\psi_P^{(1)}, \psi_P^{(2)}, \psi_P^{(3)}$  are rapidly oscillating, except in the direction of output wave). We find

for  $\psi_P^{(1)}$ , it's  $\mathbf{n}=\mathbf{n}_0$ , i.e. the wave propagates along the mirror wave direction.

for  $\psi_P^{(2)}$ ,  $\mathbf{n} = \mathbf{e}_z$ , i.e. the wave propagates along  $\mathbf{e}_z$ , the direction perpendicular to the hologram.

for  $\psi_P^{(3)}$ ,  $2\sin\theta_0 = \sin\theta$ . (Here we assume that both  $\mathbf{n}$  and  $\mathbf{n}_0$  are in the y-z plane, and  $\theta$  and  $\theta_0$  are the angles between them and  $\mathbf{e}_z$  respectively. We'll make the same assumption in part (b) and (c).) Note that if  $\theta_0 < \pi/6$ , then there exists a solution to this equation and the secondary image actually exists.

(b) Now if all angles are small (paraxial optics), then

$$\psi_P^{(3)} = \frac{ik}{2\pi} \int d\mathbf{S}' M O^* \frac{e^{ikz}}{z} e^{iky'(2\theta_0 - \theta)}, \quad \theta_0 \approx 0$$

The field at the object can be obtained by propagating back to point (0,0,-z)

$$\psi_i = rac{ik}{2\pi} \int d{f S}' Orac{e^{-ikz}}{z} e^{iky' heta_i}, \quad heta_s pprox heta_i$$

i.e.  $\psi_P(z) \propto \psi_i^*(-z)$ , we see that indeed the secondary image resides in front of the hologram and is turned inside out. See Fig. 2.

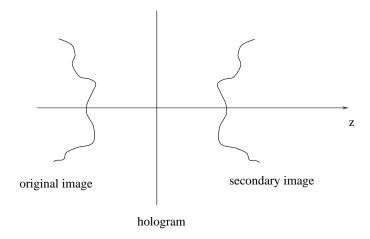


Figure 2: the secondary image

The angles in the wave scattered from the object and in the secondary wave are related by

$$-sin\theta_i = sin2\theta_0 - sin\theta_s$$

and for  $\theta_i \approx 0$ , if  $\theta_0$  is not small, then

$$\Delta heta_i = cos heta_s \Delta heta_s \Rightarrow \Delta heta_s = rac{\Delta heta_i}{cos 2 heta_0} > \Delta heta_i$$

i.e. the image is stretched.

(c) If light with wavenumber k is shined at the original angle  $\theta_0$  for reconstruction, then

$$\begin{array}{lcl} \psi_0 & \sim & MOe^{-ik_0y'sin\theta_0}e^{iky'sin\theta_0} \\ \psi_P^{(2)} & \sim & \int d\mathbf{S}' \left(\frac{\mathbf{n}+\mathbf{n}_0}{2}\right)MO\frac{e^{ikR}}{r}e^{iy'(-k_0sin\theta_0+ksin\theta_0-ksin\theta)} \end{array}$$

So again by looking for a stationary phase condition we see that the direction of propagation will be given by  $\sin \theta = \sin \theta_0 \left(1 - \frac{k_0}{k}\right)$ . (the sign of  $\theta$  is indicated in Fig. 3)

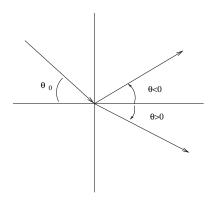


Figure 3: plane-parallel white light

$$heta_0 = \pi/4, \ k_0 = k_{green}, \ ext{and} \ sin heta = \sin heta_0 \left(1 - rac{\lambda}{\lambda_{green}}
ight).$$

(1)  $\lambda = \lambda_{green}, \, sin\theta = 0, \, \theta = 0$ 

(1) 
$$\lambda = \lambda_{green}$$
,  $sin\theta = 0$ ,  $\theta = 0$   
(2)  $\lambda = \lambda_{red}$ ,  $sin\theta = \frac{1}{\sqrt{2}} \left( 1 - \frac{\lambda_{red}}{\lambda_{green}} \right) \approx \frac{1}{\sqrt{2}} \left( 1 - \frac{700nm}{500nm} \right) \approx -0.28$ . Thus  $\theta \approx -16.4^{\circ}$ .

2. Ex 9.3 Compact disks [by Alexei Dvoretskii/99]

To record information on a CD, pits are burned in it by using a laser with  $\lambda \approx 0.5 \mu m$ . Because of diffraction it's not possible to focus the light much better than that.

So in the best case 1 bit can be written per  $0.25\mu m^2$ .

Estimate: Disc area:  $\pi \left(r_{out}^2 - r_{inner}^2\right) \approx 3[(6cm)^2 - (2cm)^2] \approx 100cm^2$ . and thus in principle the total amount of information that can be stored is  $I = \frac{100cm^2}{0.25\mu m^2} = 4 \times 10^{10} bit$ . Taking 5 per cent of this, we get  $I = 2 \times 10^9 bits$ .

Now for Encyclopedia Britannica:

1 page  $\sim 1$ kbyte  $\sim 10^4$ bit

1 volume  $\sim 10^3$  pages  $\sim 10^7 bit$ .

i.e. approximately 200 volumes could be stored on a CD.

3. Ex. 9.4 Nonlinear susceptibility for an isotropic medium [by Alexei Dvoretskii/99]

For an isotropic medium its susceptibility tensors should be symmetric in all indices. The only isotropic symmetric tensor available to us is the metric tensor  $g_{ij}$ , the Levi-Civita tensor being antisymmetric.

Therefore, all susceptibility tensors with an odd number of indices are zero  $(\chi_{ijk} = 0, \chi_{ijklm} = 0, ...)$  and the susceptibility tensors with an even number of indices must be proportional to symmetrized sums of products of metric tensors.

Thus 
$$\chi_{ij} = \chi_0 g_{ij}$$

For  $\chi_{ijkl}$ , it should be a linear combination of  $g_{ij}g_{kl}$ ;  $g_{ik}g_{jl}$ ;  $g_{il}g_{jk}$ . And since these tensors can be obtained from each other by permutations of indices their only combination symmetric w.r.t. all permutations is their sum. Hence

$$\chi_{ijkl} = rac{\chi_4}{3}(g_{ij}g_{kl} + g_{ik}g_{jl} + g_{il}g_{jk})$$

In general, a (2n)th rank susceptibility tensor will decompose for an isotropic medium into a symmetric sum of products of n metric tensors with  $\frac{(2n)!}{2^n n!}$  terms.

$$\chi_{ijklmn} = \frac{\chi_{6}}{15} \cdot \\ (g_{ij}g_{kl}g_{mn} + g_{ik}g_{jl}g_{mn} + g_{il}g_{jk}g_{mn} \\ + g_{ij}g_{km}g_{ln} + g_{ij}g_{kn}g_{lm} + g_{ik}g_{jm}g_{ln} \\ + g_{ik}g_{jn}g_{lm} + g_{il}g_{jm}g_{kn} + g_{il}g_{jn}g_{km} \\ + g_{im}g_{jk}g_{ln} + g_{im}g_{jl}g_{kn} + g_{im}g_{jn}g_{kl} \\ + g_{in}g_{jk}g_{lm} + g_{in}g_{jl}g_{km} + g_{in}g_{jm}g_{lk})$$

- 4. Ex. 9.6 Dispersion relation for an anisotropic medium [by Xinkai Wu/02]
- (a) It's known that  $P_i = \epsilon_0 \chi_{ij} E_j$  when nonlinear effects are unimportant. Hence we have

$$D_i = \epsilon_0 E_i + P_i = \epsilon_0 E_i + \epsilon_0 \chi_{ij} E_j = \epsilon_0 \epsilon_{ij} E_j$$

if we define  $\epsilon_{ij} \equiv \delta_{ij} + \chi_{ij}$ .

(b) This is very straightforward (recall that by definition  $\mathbf{D} = \epsilon_0 \mathbf{E} + \mathbf{P}$ ).

$$\nabla \cdot \mathbf{D} = \epsilon_0 \nabla \cdot \mathbf{E} + \nabla \cdot \mathbf{P}$$
using the first identity in eq. (9.15)
$$= \epsilon_0 \left( -\frac{\nabla \cdot \mathbf{P}}{\epsilon_0} \right) + \nabla \cdot \mathbf{P}$$

$$= 0$$

Also the last identity in eq. (9.15) reads

$$\nabla \times \mathbf{B} = \mu_0 \left( \frac{\partial \mathbf{P}}{\partial t} + \epsilon_0 \frac{\partial \mathbf{E}}{\partial t} \right)$$
$$= \mu_0 \frac{\partial}{\partial t} (\mathbf{P} + \epsilon_0 \mathbf{E})$$
$$= \mu_0 \frac{\partial \mathbf{D}}{\partial t}$$

The other two sourceless equations remain the same as in eq. (9.15)

(c) Taking the curl of the third identity in eq. (9.37), we get

$$\nabla \times (\nabla \times \mathbf{E}) = \nabla (\nabla \cdot \mathbf{E}) - \nabla^{2} \mathbf{E}$$

$$= \nabla \times \left( -\frac{\partial \mathbf{B}}{\partial t} \right) = -\frac{\partial}{\partial t} (\nabla \times \mathbf{B})$$
using the fourth identity in eq. (9.37)
$$= -\mu_{0} \frac{\partial^{2} \mathbf{D}}{\partial t^{2}}$$
using the fact that  $\mathbf{D} = \epsilon_{0} \epsilon \cdot \mathbf{E}$ 

$$= -\mu_{0} \epsilon_{0} \epsilon \cdot \frac{\partial^{2} \mathbf{E}}{\partial t^{2}}$$

$$= -\epsilon \cdot \frac{\partial^{2} \mathbf{E}}{c^{2} \partial t^{2}}$$

(d) Consider the plane wave  $E_j(\mathbf{x},t) = E_j e^{i(\mathbf{k}\cdot\mathbf{x}-\omega t)}$ , where  $E_j$  is a constant amplitude. Plugging this plane waveform into eq. (9.38), one readily get

$$L_{ij}E_{j}=0, \;\; ext{where} \; L_{ij}=k_{i}k_{j}-k^{2}\delta_{ij}+rac{\omega^{2}}{c^{2}}\epsilon_{ij}$$

(e) Consider waves propagating in the x-z plane,  $\mathbf{k} = k(sin\theta\mathbf{e}_x + cos\theta\mathbf{e}_z)$ . Then it's straightforward (and we omit the details here) to find that  $||L_{ij}|| = k^2 \cdot [$ the matrix given in eq. (9.41)], where  $k = n\frac{\omega}{c}$ . Computing the determinant of this matrix and pulling out some factors, one find that the dispersion relation (9.40) reduces to

$$\left(\frac{1}{n^2} - \frac{1}{n_o^2}\right) \left(\frac{1}{n^2} - \frac{\cos^2 \theta}{n_o^2} - \frac{\sin^2 \theta}{n_e^2}\right) = 0$$

(f) Solving the dispersion relation we found above, we find two solutions

(i) 
$$n = n_o$$
.

In this case, solving for the eigenvectors, one finds there's no constraint on  $E_y$ , while  $E_x$  and  $E_z$  must satisfy

$$egin{aligned} sin^2 heta E_x + sin heta cos heta E_z &= 0 \ sin heta cos heta E_x + \left[\left(rac{n_e}{n_o}
ight)^2 - sin^2 heta
ight] E_z &= 0 \end{aligned}$$

the determinant of the coefficient matrix for the above 2 equations is  $sin^2\theta\left[\left(\frac{n_e}{n_o}\right)^2-1\right]$ , which is nonzero if we consider  $sin\theta\neq 0$  and  $n_e\neq n_o$ . Hence  $E_x=0, E_z=0$ . And E is along the y-direction.

This is the polarization described in eq. (9.30) and associated text.

(ii) 
$$\frac{1}{n} = \sqrt{\frac{\cos^2\theta}{n_o^2} + \frac{\sin^2\theta}{n_e^2}}$$

(ii)  $\frac{1}{n} = \sqrt{\frac{\cos^2\theta}{n_o^2} + \frac{\sin^2\theta}{n_e^2}}$ . Solving for the eigenvectors, one finds that (again assuming  $\sin\theta \neq 0$  and  $n_e \neq n_o$ )  $E_y = 0$ , and

$$\left(rac{n_o}{n_e}
ight)^2 sin heta E_x + cos heta E_z = 0$$

i.e.

$$\frac{E_x}{E_z} = -\left(\frac{n_e}{n_o}\right)^2 \cot\theta$$

This is the polarization described in eq. (9.31) and associated text.