## Solution for Chapter 5

(compiled by Xinkai Wu, revised by Kip Thorne)

1.

Ex. 5.3 Wiener's Optimal Filter [by Alexei Dvoretskii]

(a). Let's first show  $\overline{N(t)}=0$ . One can do this with the help of ergodicity:  $\overline{N(t)}=< N(t)>=\int_{-\infty}^{+\infty}K(t-t')< y(t')>dt'=\int_{-\infty}^{+\infty}K(t-t')\overline{y(t')}dt'=0$ . One can also prove this directly by integrating N(t) over time, whose detail we omit here.

Knowing N(t) = 0, we readily get

$$\overline{N^2(t)} = \sigma_N^2 = \int_0^{+\infty} S_N(f) df = \int_0^{+\infty} |\tilde{K}(f)|^2 S_y(f) df.$$

By ergodicity  $\overline{N^2(t)} = \langle N^2(t) \rangle$ . Now let's show  $\langle N^2(t) \rangle = \langle N^2 \rangle$ .

$$< N^{2}(t) > = < \int_{-\infty}^{+\infty} K(t - t') y(t') dt' \int_{-\infty}^{+\infty} K(t - t'') y(t'') dt'' >$$

$$= \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} dt' dt'' K(t - t') K(t - t'') < y(t') y(t'') >$$

$$= \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} dt' dt'' K(t - t') K(t - t'') f(t' - t'')$$

$$let \qquad \tilde{t}' = t - t', \ \tilde{t}'' = t - t''$$

$$= \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} d\tilde{t}' d\tilde{t}'' K(\tilde{t}') K(\tilde{t}'') f(\tilde{t}'' - \tilde{t}')$$

$$= \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} dt' dt'' K(t') K(t'') f(t'' - t')$$

In the above equations we use f(t'-t'') to denote  $\langle y(t')y(t'') \rangle$  because it's a stationary process.

On the other hand, we have

$$egin{array}{lcl} < N^2 > &=& < \int_{-\infty}^{+\infty} K(t') y(t') dt' \int_{-\infty}^{+\infty} K(t'') y(t'') dt'' > \ &=& \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} dt' dt'' K(t') K(t'') f(t'-t'') \end{array}$$

Using the trivial fact that f(t'-t'')=f(t''-t'), we see that  $< N^2>=< N^2(t)>=\overline{N^2(t)}=\int_0^{+\infty}|\tilde{K}(f)|^2S_y(f)df$ .

[A much simpler proof by Kip:  $N(t) = \int_{-\infty}^{+\infty} K(t'-t)y(t')dt' = \int_{-\infty}^{+\infty} K(t'')y(t''+t)dt''$  where we've made a change of variable  $t'' \equiv t'-t$ . But y is a stationary random process, so its statistical properties are independent of the origin of time:  $p_1(y,t) = p_1(y,0)$ . Thus, the statistical properties of N(t) must be the same as those of N. In particular,  $N^2(t) > = N^2 >$ 

(b) Using Parseval's theorem, and the fact K(t), s(t) are both real (thus  $\tilde{K}(-f) = \tilde{K}^*(f), \tilde{s}(-f) = \tilde{s}^*(f)$ ), we have

$$S = \int_{-\infty}^{+\infty} K(t)s(t)dt = \frac{1}{2} \int_{-\infty}^{+\infty} (\tilde{K}(f)\tilde{s}^*(f) + c.c)df = \int_{0}^{+\infty} (\tilde{K}(f)\tilde{s}^*(f) + c.c)df$$

We have

$$\frac{S}{\langle N^2 \rangle^{1/2}} = \frac{\int_0^{+\infty} (\tilde{K}(f)\tilde{s}^*(f) + c.c)df}{\left[ \int_0^{+\infty} |\tilde{K}(f)|^2 S_y(f)df \right]^{1/2}}$$

Taking a small variation  $\tilde{K}(f) o \tilde{K}(f) + \delta \tilde{K}(f),$  one readily gets

$$\delta\left(rac{S}{< N^2>^{1/2}}
ight) = rac{S}{< N^2>^{1/2}} \left[ \int_0^{+\infty} df \delta ilde{K}(f) \left(rac{ ilde{s}(f)}{S} - rac{ ilde{K}(f)S_y(f)}{2< N^2>}
ight) + c.c.
ight]$$

We see that for  $\delta\left(\frac{S}{< N^2 > ^{1/2}}\right)$  to vanish for any  $\delta \tilde{K}(f)$ , we must have

$$ilde{K}(f) = const imes rac{ ilde{s}(f)}{S_u(f)}$$

To show that this choice of K(f) actually delivers a maximum to the ratio  $\frac{S}{\langle N^2 \rangle^{1/2}}$ , one could do a second variation calculation. Our physical intuition should make it obvious, though, because a filter with such a kernel favors such frequencies for which the signal to noise ratio is high and suppresses theose for which the opposite is true.

2.

Ex. 5.4 Alan Variance of Clocks [by Xinkai Wu]

(a) We just need to find the relation between the Fourier transform of different random processes.

$$\begin{split} \tilde{\Phi}_{\tau}(f) &= \frac{1}{\sqrt{2}\bar{\omega}\tau}\tilde{\phi}(f)\left(e^{-i2\pi f\cdot 2\tau} - 2e^{-i2\pi f\tau} + 1\right) \\ &= \frac{\sqrt{2}}{\bar{\omega}\tau}\tilde{\phi}(f)e^{-i2\pi f\tau}(\cos 2\pi f\tau - 1) \end{split}$$

Also, since  $\phi(t)$  is obtained by integrating  $\omega(t)$  once, we have

$$\tilde{\phi}(f) = \frac{-1}{i2\pi f}\tilde{\omega}(f)$$

Combining the above two results, and using the fact that the spectral density is basically given by the modulus squared Fourier transform, we find

$$S_{\Phi_{\tau}(f)} = \frac{2}{\bar{\omega}^2} \left[ \frac{\cos 2\pi f \tau - 1}{2\pi f \tau} \right]^2 S_{\omega}(f)$$

$$\propto f^2 S_{\omega}(f) \text{ for } f << 1/2\pi \tau$$

$$\propto f^{-2} S_{\omega}(f) \text{ for } f >> 1/2\pi \tau$$

[An alternative way of finding  $S_{\Phi_{\tau}}(f)$ : As discussed in Section 5.5 of the text,  $\Phi_{\tau}(t)$  can be regarded as obtained from  $\omega(t')$  using a filter K(t-t'). Then  $S_{\Phi_{\tau}}(f) = \left| \tilde{K}(f) \right|^2 S_{\omega}(f)$ . To find  $\left| \tilde{K}(f) \right|^2$  we feed  $\omega(t') = \exp[i2\pi ft']$  into our system. We find, using eqn (5.109) and (5.110),  $\Phi_{\tau}(t) = \frac{\exp[i2\pi f(t+\tau)]\sqrt{2}(\cos 2\pi f\tau-1)}{\bar{\omega}\tau i2\pi f}$ , whose modulus square gives  $\left| \tilde{K}(f) \right|^2 = \frac{2}{\bar{\omega}^2} \left[ \frac{\cos 2\pi f\tau-1}{2\pi f\tau} \right]^2$ . Thus we arrive at eqn (5.111)]

(b) Using the expression for  $S_{\Phi_{\tau}}(f)$  obtained in the previous part, and making the change of variable  $z \equiv 2\pi f \tau$  in the integral, one finds

$$\sigma_{\tau} = \left[ \alpha \frac{S_{\omega}(1/2\tau)}{\bar{\omega}^2} \frac{1}{2\tau} \right]^{1/2}$$

$$where \ \alpha = \int_0^{+\infty} dz \frac{2}{\pi} \left[ \frac{\cos z - 1}{z} \right]^2 \frac{S_{\omega}(z/2\pi\tau)}{S_{\omega}(1/2\tau)}$$

As one can verify,  $\frac{2}{\pi} \left[ \frac{\cos z - 1}{z} \right]^2$  integrates to one, and has a profile shown in

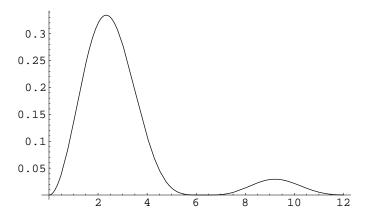


Figure 1: Ex. 5.4 (b)

Fig.1, which isn't too much different from a delta function located at z=2.4 (which isn't too much away from  $z=\pi$ ). Thus we see  $\alpha$  is a dimensionless number of order unifty and has very weak  $\tau$ -dependence which we can ignore.

(c) If  $\omega$  has a white-noise spectrum  $S_{\omega}(1/2\tau) \sim const$ , then  $\sigma_{\tau} \propto 1/\sqrt{\tau}$ , and the clock stability is better for long averaging time; if  $\omega$  has a flicker-noise spectrum  $S_{\omega}(1/2\tau) \propto \tau$ , then  $\sigma_{\tau}$  is independent of averaging time; if  $\omega$  has a random-walk spectrum  $S_{\omega}(1/2\tau) \propto \tau^2$ , then  $\sigma_{\tau} \propto \sqrt{\tau}$ , and the clock stability is better for short averaging time.

Ex. 5.5 Cosmological Density Fluctuations [by Roger Blandford]

(a) It's quite straightforward to show this,

$$\begin{split} \xi(\mathbf{r}) &= \langle \delta(\mathbf{x})\delta(\mathbf{x}+\mathbf{r}) > = \lim_{V \to \infty} \frac{1}{V} \int_{V} d\mathbf{x} \delta(\mathbf{x})\delta(\mathbf{x}+\mathbf{r}) \\ &= \lim_{V \to \infty} \frac{1}{V} \int_{V} d\mathbf{x} \frac{1}{(2\pi)^{3}} \int d\mathbf{k} e^{-i\mathbf{k}\cdot\mathbf{x}} \tilde{\delta}_{V}(\mathbf{k}) \frac{1}{(2\pi)^{3}} \int d\mathbf{k}' e^{-i\mathbf{k}'\cdot(\mathbf{x}+\mathbf{r})} \tilde{\delta}_{V}(\mathbf{k}') \\ &= [\operatorname{performing} \ the \ integral \ in \ \mathbf{x} \ gives \ a \ delta \ function \ in \ (\mathbf{k}+\mathbf{k}')] \\ &= \lim_{V \to \infty} \frac{1}{V} \frac{1}{(2\pi)^{3}} \int d\mathbf{k} \ e^{i\mathbf{k}\cdot\mathbf{r}} \tilde{\delta}_{V}(\mathbf{k}) \tilde{\delta}_{V}(-\mathbf{k}) \\ &= [\delta(\mathbf{x}) \ is \ real \Rightarrow \tilde{\delta}_{V}(-\mathbf{k}) = \tilde{\delta}_{V}^{*}(\mathbf{k}); \ also \ let \ \mathbf{k} \to -\mathbf{k} \ in \ the \ integral] \\ &= \int \frac{d\mathbf{k}}{(2\pi)^{3}} e^{-i\mathbf{k}\cdot\mathbf{r}} lim_{V \to \infty} \frac{|\tilde{\delta}_{V}(\mathbf{k})|^{2}}{V} = \int \frac{d\mathbf{k}}{(2\pi)^{3}} e^{-i\mathbf{k}\cdot\mathbf{r}} P(\mathbf{k}) \end{split}$$

The universe is isotropic, namely,  $\delta(\mathbf{x}) = \delta(|\mathbf{x}|) \Rightarrow \tilde{\delta}_V(\mathbf{k}) = \tilde{\delta}_V(k) \Rightarrow P(\mathbf{k}) = P(k)$ . And we can perform the momentum integral in spherical coordinates. Using the fact  $\int_0^{\pi} d\theta sin\theta exp(-ikrcos\theta) = 2sinc(kr)$ , one finds

$$\xi(\mathbf{r}) = \int_0^\infty \frac{dk}{2\pi^2} k^2 sinc(kr) P(k)$$

(b) The mass measured within a sphere of radius R is given by

$$egin{array}{lcl} \delta_R({f x}) &=& rac{3}{4\pi R^3} \int_{r < R} d{f r} \delta({f x} + {f r}) \ &=& \int_V d{f r} \delta({f x} + {f r}) K({f r}) \ where \ K({f r}) &=& rac{3}{4\pi R^3} \ for \ \ r < R, \ and \ vanishes \ elsewhere. \end{array}$$

One can regard  $K(\mathbf{r})$  as a filter.

By derivation similar to that in the previous part, one finds the variance

$$\sigma^2 = \langle \delta_R^2(\mathbf{x}) \rangle = \lim_{V \to \infty} \int \frac{1}{V} \frac{d\mathbf{k}}{(2\pi)^3} |\tilde{\delta}_R(\mathbf{k})|^2$$

By Parseval's theorem,  $|\tilde{\delta}_R(\mathbf{k})|^2 = |\tilde{\delta}_V(\mathbf{k})|^2 |\tilde{K}(\mathbf{k})|^2$ . Using spherical coordinates, one finds  $\tilde{K}(\mathbf{k}) = W(kR)$ , where  $W(x) \equiv \frac{3(sincx-cosx)}{x^2}$ . Combining the above results, we get

$$\sigma^2 = \int_0^\infty \frac{dk}{2\pi^2} k^2 P(k) W^2(kR)$$

3.

Ex. 5.7 Noise in a L-C-R Circuit [by Alexei Dvoretskii]

- (a)  $V_{\alpha\beta} = R\dot{q} F'$ , where  $\dot{q}$  is the current in the circuit and q is the charge on the capacitor. When we disconnect the resistor R from the rest of the circuit,  $\dot{q} = 0$ , and thus the spectral density of  $V_{\alpha\beta}$  is the same as that of F', namely  $S_{\alpha\beta}(f) = 4RkT$ .
- (b) Now let's place the resistor back into the circuit, and  $\dot{q}$  is no longer zero. Using  $L\ddot{q} + C^{-1}q = -R\dot{q} + F'$  we can find the Fourier transform of q (and thus that of  $V_{\alpha\beta}$ ) in terms of the Fourier transform of F'. So we get

$$S_{\alpha\beta}(f) = \left| \frac{2\pi i f R}{C^{-1} - L(2\pi f)^2 + 2\pi i f R} - 1 \right|^2 S_{F'}(f)$$

$$[using \ S_{F'}(f) = 4RkT, \ (2\pi f_0)^2 = \frac{1}{LC}]$$

$$\Rightarrow S_{\alpha\beta}(f) = 4RkT \left\{ 1 - \frac{(2\pi f)^2 R^2}{L^2 \left[ (2\pi f_0)^2 - (2\pi f)^2 \right]^2 + (2\pi f)^2 R^2} \right\}$$

(c)  $V_{\alpha\gamma} = \frac{q}{C}$ , thus

$$S_{lpha\gamma}(f) = rac{1}{C^2} rac{1}{\left|C^{-1} - L(2\pi f)^2 + 2\pi i f R
ight|^2} 4RkT$$

(d)  $V_{\beta\gamma} = -L\ddot{q}$ , thus

$$S_{eta\gamma}(f) = L^2 (2\pi f)^4 rac{1}{\left|C^{-1} - L(2\pi f)^2 + 2\pi i f R
ight|^2} 4RkT$$

(e) Similar to Ex 5.3, the number U is given by

$$U = \int_{-\infty}^{+\infty} K(t) V_{lphaeta}(t) dt,$$
 with  $K(t) = rac{1}{ au} \ for \ - au < t < 0, \ and \ 0 \ elsewhere$ 

We also introduce the random process  $U(t) \equiv \int_{-\infty}^{+\infty} K(t-t') V_{\alpha\beta}(t') dt'$ . Then easily seen  $\bar{U} = 0$ , and

$$(\Delta U)^2 = \overline{U(t)^2} = \int_0^\infty \left| \tilde{K}(f) \right|^2 S_{\alpha\beta}(f) df$$

Using the fact that  $\left| \tilde{K}(f) \right|^2 = \left[ \frac{\sin(\pi f \tau)}{\pi f \tau} \right]^2$ , [Again, just like in Ex. 5.4, an alternative way of finding  $\left| \tilde{K}(f) \right|^2$  is: note  $U(t) = \frac{1}{\tau} \int_t^{t+\tau} V_{\alpha\beta}(t') dt'$ , thus  $\left| \tilde{K}(f) \right|^2 = \left| \frac{1}{\tau} \int_t^{t+\tau} exp[i2\pi f t'] dt' \right|^2 = \left[ \frac{\sin(\pi f \tau)}{\pi f \tau} \right]^2$ ] and plugging in the expression

sion for  $S_{\alpha\beta}(f)$  we found in part (b), we get (define  $x \equiv \pi f \tau$ )

$$\begin{array}{lll} (\Delta U)^2 & = & \displaystyle \frac{4RkT}{\pi\tau} \left\{ \int_0^\infty \frac{\sin^2 x}{x^2} dx - \int_0^\infty \frac{\sin^2 x}{\left[x^2 - (\pi\tau f_0)^2\right]^2 \left(\frac{2L}{\tau R}\right)^2 + x^2} dx \right\} \\ & = & \displaystyle \frac{2RkT}{\tau} \left\{ 1 - \frac{RC}{\tau} \left[ 1 - e^{-\frac{R\tau}{2L}} \cos \left(\tau \sqrt{\frac{1}{LC} - \left(\frac{R}{2L}\right)^2}\right) \right] \right\} \end{array}$$

Note in the above expression  $\frac{RC}{\tau} = \frac{1}{2Q\pi\tau f_0}$ , where  $Q \equiv \frac{2\pi f_0 L}{R}$  is the quality factor of the circuit. In our case, Q >> 1, and  $\pi\tau f_0 >> 1$ , thus  $\frac{RC}{\tau} << 1$ . So approximately  $(\Delta U)^2 \approx \frac{2RkT}{\tau}$ .

4.

Ex. 5.8 Thermal Noise in a Sapphire Crystal [by Alexei Dvoretskii]

(a) Assume we are in the classical regime kT>>hf. Then  $S_{F'}(f)=4RkT$ , with  $R=ReZ=Re\left(\frac{\tilde{F}'(f)}{-i2\pi f\tilde{x}(f)}\right)=\frac{2m}{\tau_*}$ . Thus  $S_{F'}(f)=\frac{8mkT}{\tau_*}$ . And

$$S_x(f) = \frac{S_{F'}(f)}{\left| m(\omega^2 - (2\pi f)^2) + \frac{4\pi i f m}{\tau_*} \right|^2}$$
$$= \frac{\frac{8mkT}{\tau_*}}{\left| m(\omega^2 - (2\pi f)^2) + \frac{4\pi i f m}{\tau_*} \right|^2}$$

(b)  $m(\ddot{x} + \frac{2}{\tau_*}\dot{x} + \omega^2 x) = Re\left(\sqrt{2}F_s e^{i\omega t}\right)$   $let \ x = Re\left(x_0 e^{i\omega t}\right)$   $we \ get \ x_0 = \frac{-i\tau_*}{\sqrt{2}m\omega}F_s$   $thus \ x(t) = \frac{\tau_* F_s}{\sqrt{2}m\omega}sin\omega t$ 

(c) After filtering, signal  $x(t)=\sqrt{2}x_ssin(\omega t)$  with  $x_s=\left|\tilde{K}(\omega/2\pi)\right|\frac{\tau_*F_s}{2m\omega}$ ; noise  $\sigma_x=\left|\tilde{K}(\omega/2\pi)\right|\sqrt{S_x(f)\Delta f}$ , where  $S_x(f=\omega/2\pi)=\frac{kT\tau_*}{2\pi^2f^2m}$ . So we see, if  $x_s=\sigma_x$ , then  $F_s=\frac{2m\omega}{\tau_*}\sqrt{\frac{kT\tau_*\Delta f}{2\pi^2f^2m}}=\left(\frac{8mkT}{\tau_*\hat{\tau}}\right)^{1/2}$ . (d)

$$h_s \sim \frac{2}{m\omega^2 l} \left(\frac{2mkT}{\tau_*\hat{ au}}\right)^{1/2} = 2\sqrt{2} \left(\frac{kT}{ml^2\omega^2}\right)^{1/2} \frac{1}{(\omega\hat{ au})^{1/2}(\pi Q)^{1/2}} \sim 1.5 \times 10^{-26}$$

Ex. 5.9 Solution of Fokker-Planck Equation for an Oscillator [by Alexander Putilin]

(a) Equation of motion:  $\ddot{x} + \frac{2}{\tau_*}\dot{x} + \omega^2 x = F'/m$ . Substitute:  $x = X_1(t)cos\omega t + X_2(t)sin\omega t$  into this e.o.m., we get

$$cos\omega t\left(\ddot{X}_1+\frac{2}{\tau_*}\dot{X}_1+2\omega\dot{X}_2+\frac{2\omega}{\tau_*}X_2\right)+sin\omega t\left(\ddot{X}_2+\frac{2}{\tau_*}\dot{X}_2-2\omega\dot{X}_1-\frac{2\omega}{\tau_*}X_1\right)=\frac{F'}{m}$$

 $\{X_1, X_2\}$  evolve on a time scale  $\tau_*$  means that  $\dot{X}_i \sim \frac{X_i}{\tau_*}$ ,  $\ddot{X}_i \sim \frac{X_i}{\tau_*^2}$ , with i=1,2. And since  $\tau_* >> \omega^{-1}$ , we can neglect the first two terms in the brackets. Thus the Langevin equation reduces to:

$$-2\omega(\dot{X}_1 + X_1/\tau_*)sin\omega t + 2\omega(\dot{X}_2 + X_2/\tau_*)cos\omega t = F'/m$$

(b) Multiply the Langevin equation by  $cos\omega t$  and integrate from t=0 to  $t=\Delta t$  where  $\omega^{-1}<<\Delta t<<\tau_*$ :

$$\begin{split} &2\omega\int_{0}^{\Delta t}\left(\dot{X}_{2}+\frac{1}{\tau_{*}}X_{2}\right)cos^{2}\omega tdt-2\omega\int_{0}^{\Delta t}\left(\dot{X}_{1}+\frac{2}{\tau_{*}}X_{1}\right)cos\omega tsin\omega tdt\\ &=\frac{1}{m}\int_{0}^{\Delta t}F'(t)cos\omega tdt \end{split}$$

 $\cos^2 \omega t$  and  $\cos \omega t \sin \omega t$  are fast oscillating functions, so we use their average value inside the integrals, i.e. change  $\cos^2 \omega t \to \overline{\cos^2 \omega t} = \frac{1}{2}$ ,  $\cos \omega t \sin \omega t \to \overline{\cos \omega t \sin \omega t} = 0$ . Thus we get

$$\omega \int_0^{\Delta t} \left( \dot{X}_2 + \frac{1}{\tau_*} X_2 \right) dt \approx \omega \left( \Delta X_2 + \frac{1}{\tau_*} X_2 \Delta t \right) = \frac{1}{m} \int_0^{\Delta t} F'(t) cos\omega t dt$$
i.e.  $\Delta X_2 = -\frac{1}{\tau_*} X_2 \Delta t + \frac{1}{m\omega} \int_0^{\Delta t} F'(t) cos\omega t dt$ 

Similarly, multiplying the Langevin equation by  $sin\omega t$  and integrating, we get

$$\Delta X_1 = -\frac{1}{\tau_*} X_1 \Delta t - \frac{1}{m\omega} \int_0^{\Delta t} F'(t) sin\omega t dt$$

(c) Taking ensemble average:

$$\begin{split} \overline{\Delta X_1} &= \frac{-1}{\tau_*} X_1 \Delta t - \frac{1}{m\omega} \int_0^{\Delta t} \overline{F'(t)} sin\omega t dt \\ \overline{F'(t)} &= 0 \Rightarrow A_1 = \frac{\overline{\Delta X_1}}{\Delta t} = \frac{-X_1}{\tau_*} \\ \text{similarly we have } A_2 &= \frac{\overline{\Delta X_2}}{\Delta t} = \frac{-X_2}{\tau_*} \\ \text{thus } A_j &= \frac{-X_j}{\tau_*} \end{split}$$

now consider 
$$B_{jk} = \frac{\overline{\Delta X_j \Delta X_k}}{\Delta t}$$

$$\begin{split} B_{11} &= \frac{\overline{(\Delta X_1)^2}}{\Delta t} &= \frac{1}{m^2 \omega^2 \Delta t} \int_0^{\Delta t} dt_1 dt_2 sin\omega t_1 sin\omega t_2 < F'(t_1) F'(t_2) > + O\left[(\Delta t)\right] \\ &= \frac{1}{m^2 \omega^2 \Delta t} \int_0^{\Delta t} dt_1 dt_2 sin\omega t_1 sin\omega t_2 C_{F'}(t_1 - t_2) \\ &= \frac{1}{m^2 \omega^2 \Delta t} \int_0^{\Delta t} dt_1 dt_2 sin\omega t_1 sin\omega t_2 \int_0^{+\infty} S_{F'}(f) cos[2\pi f(t_1 - t_2)] df \\ &= \frac{1}{m^2 \omega^2 \Delta t} \int_0^{\Delta t} dt_1 dt_2 sin\omega t_1 sin\omega t_2 \int_0^{+\infty} S_{F'}(f) dt_1 dt_2 sin\omega t_1 sin\omega t_2 \int_0^{+\infty} cos[2\pi f(t_1 - t_2)] df \\ &= \frac{8kT}{m\omega^2 \tau_* \Delta t} \int_0^{\Delta t} dt_1 dt_2 sin\omega t_1 sin\omega t_2 \int_0^{+\infty} cos[2\pi f(t_1 - t_2)] df \\ &= \frac{4kT}{m\omega^2 \tau_* \Delta t} \int_0^{\Delta t} dt_1 sin^2 \omega t_1 \approx \frac{4kT}{m\omega^2 \tau_* \Delta t} \cdot \frac{\Delta t}{2} = \frac{2kT}{m\omega^2 \tau_*} \end{split}$$

Similar calculation gives:

$$\begin{split} B_{22} &= \frac{4kT}{m\omega^2\tau_*\Delta t} \int_0^{\Delta t} dt cos^2\omega t \approx \frac{2kT}{m\omega^2\tau_*} \\ B_{12} &= B_{21} = \frac{4kT}{m\omega^2\tau_*\Delta t} \int_0^{\Delta t} dt cos\omega t sin\omega t \approx 0 \end{split}$$

Thus we conclude

$$B_{jk} = \frac{2kT}{m\omega^2\tau_*}\delta_{jk}$$

(d) After plugging the expressions for  $A_j$  and  $B_{jk}$  into the Fokker-Planck equation for  $P_2 = P_2(X_j, t|X_j^{(0)})$ , we get

$$\frac{\partial}{\partial t}P_2 = \frac{1}{\tau_*} \frac{\partial}{\partial X_j} (X_j P_2) + \frac{kT}{m\omega^2 \tau_*} \nabla^2 P_2$$

If you don't have so much time, you can just verify that Eq. (5.171) is a solution of this equation, and that it has the required form at t = 0:  $P_2(\mathbf{X}, 0) = \delta(\mathbf{X} - \mathbf{X}^{(0)})$ . If you do have some time, you could derive the solution as follows.

Make Fourier transform:

$$P_2(\mathbf{X},t) = \int \frac{d\mathbf{K}}{(2\pi)^2} e^{-i\mathbf{K}\cdot\mathbf{X}} \tilde{P}_2(\mathbf{K},t)$$

The Fokker-Planck equation becomes

$$\frac{\partial}{\partial t}\tilde{P}_2 = \frac{-1}{\tau_*} K_j \frac{\partial}{\partial K_i} \tilde{P}_2 - \frac{kT}{m\omega^2 \tau_*} \mathbf{K}^2 \tilde{P}_2$$

Make a change of variable,  $\mathbf{K} \to \mathbf{K}' \equiv e^{-t/\tau_*} \mathbf{K}$ . Then

$$K'_{j} \frac{\partial}{\partial K'_{j}} = K_{j} \frac{\partial}{\partial K_{j}}, \quad \frac{\partial}{\partial t}|_{\mathbf{K}'} = \frac{\partial}{\partial t}|_{\mathbf{K}} + \frac{1}{\tau_{*}} K'_{j} \frac{\partial}{\partial K'_{j}}$$

And the F-P equation reduces to

$$\frac{\partial}{\partial t}\tilde{P}_2(\mathbf{K}',t) = -\frac{kT}{m\omega^2\tau_*}\mathbf{K'}^2 e^{2t/\tau_*}\tilde{P}_2(\mathbf{K}',t)$$

which is easily integrated to give

$$\begin{split} \tilde{P}_2(\mathbf{K}',t) &= \tilde{P}_2(\mathbf{K}',0) exp \left[ -\frac{kT}{m\omega^2 \tau_*} \mathbf{K}'^2 \int_0^t dt' e^{2t'/\tau_*} \right] \\ \text{namely, } \tilde{P}_2(\mathbf{K},t) &= \tilde{P}_2(e^{-t/\tau_*} \mathbf{K},0) exp \left[ -\frac{kT}{2m\omega^2} \left( 1 - e^{-2t/\tau_*} \right) \mathbf{K}^2 \right] \end{split}$$

Using the initial condition  $P_2(\mathbf{X}, 0) = \delta(\mathbf{X} - \mathbf{X}^{(0)})$ , i.e.  $\tilde{P}_2(\mathbf{K}', 0) = e^{i\mathbf{K}' \cdot \mathbf{X}^{(0)}}$ , we finally find

$$\begin{split} \tilde{P}_2(\mathbf{K},t) &= e^{-\frac{1}{2}\sigma^2\mathbf{K}^2 + i\mathbf{K}\cdot\bar{\mathbf{X}}} \\ \text{where } \sigma^2 &\equiv \frac{kT}{m\omega^2} \left(1 - e^{-2t/\tau_*}\right) \text{ and } \bar{\mathbf{X}} \equiv \mathbf{X}^{(0)} e^{-t/\tau_*} \end{split}$$

Making an inverse Fourier transform (which is easy because  $\tilde{P}_2(\mathbf{K}, t)$  is Gaussian), we find

$$P_2(\mathbf{X},t) = rac{1}{2\pi\sigma^2} exp \left[ - \left. rac{\left| \mathbf{X} - ar{\mathbf{X}} 
ight|^2}{2\sigma^2} 
ight]$$

(e) At small times  $t << \tau_*$ , the probability distribution  $P_2(\mathbf{X},t)$  is concentrated around the initial value  $\mathbf{X} = \mathbf{X}^{(0)}$ . The variance of the distribution  $\sqrt{\sigma^2} \propto \sqrt{t}$ , which is the usual relation for random-walk processes. When t becomes large  $(t >> \tau_*)$ , the system "forgets" about its initial configuration and evolves into the thermal equilibrium state with  $\bar{\mathbf{X}} = 0$  and  $\sigma^2 = \frac{kT}{m\omega^2}$ .

When the signal acts for a short time  $\hat{\tau} << \tau_*$ , the thermal noise is  $\overline{\sigma^2} = \frac{1}{\hat{\tau}} \int_0^{\hat{\tau}} dt \sigma^2(t) = \frac{kT}{m\omega^2} \frac{\hat{\tau}}{\tau_*}$ , while the amplitude of the signal is  $x_s = \frac{\tau_* F_s}{2m\omega}$ , where  $F_s$  is the strength of the signal force (see part (b) of the previous exercise). Then the minimal detectable force is given by the relation  $x_s = \sqrt{\overline{\sigma^2}}$ , namely,  $F_s = \sqrt{\frac{4mkT\hat{\tau}}{\tau_*^3}}$ . In the opposite limit  $\hat{\tau} >> \tau_*$  (Ex 5.8),  $F_s = \sqrt{\frac{8mkT}{\tau_*\hat{\tau}}}$ .